

# Physics Models and Simulations of Sensors

Craig Lage

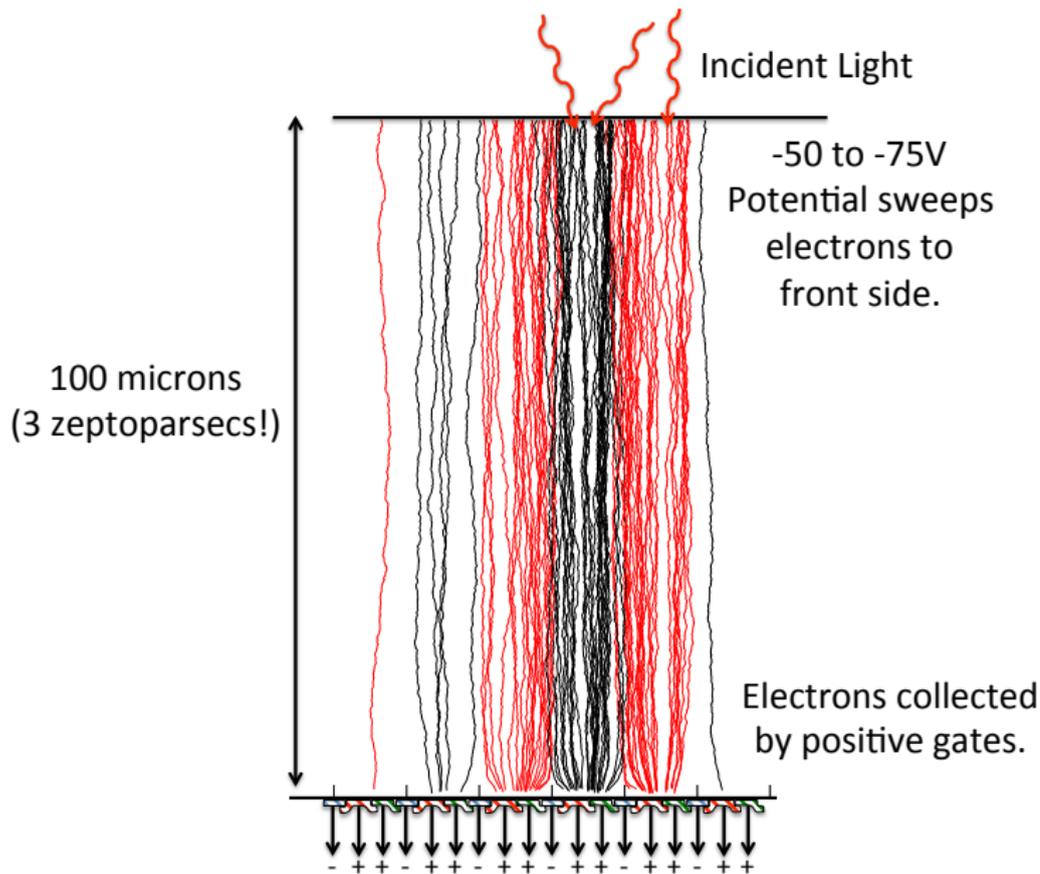
December 15, 2015

Acknowledgements:

Tony Tyson, Andrew Bradshaw, Kirk Gilmore, Perry Gee

- Introduction
- Why build a physics-based model of the sensor?
- Simulating the sensor.
- Simulation successes:
  - Edge effects.
  - Diffusion effects.
  - Brighter-Fatter effect
    - Direct measurements
    - Correlation measurements
- Conclusions and next steps.

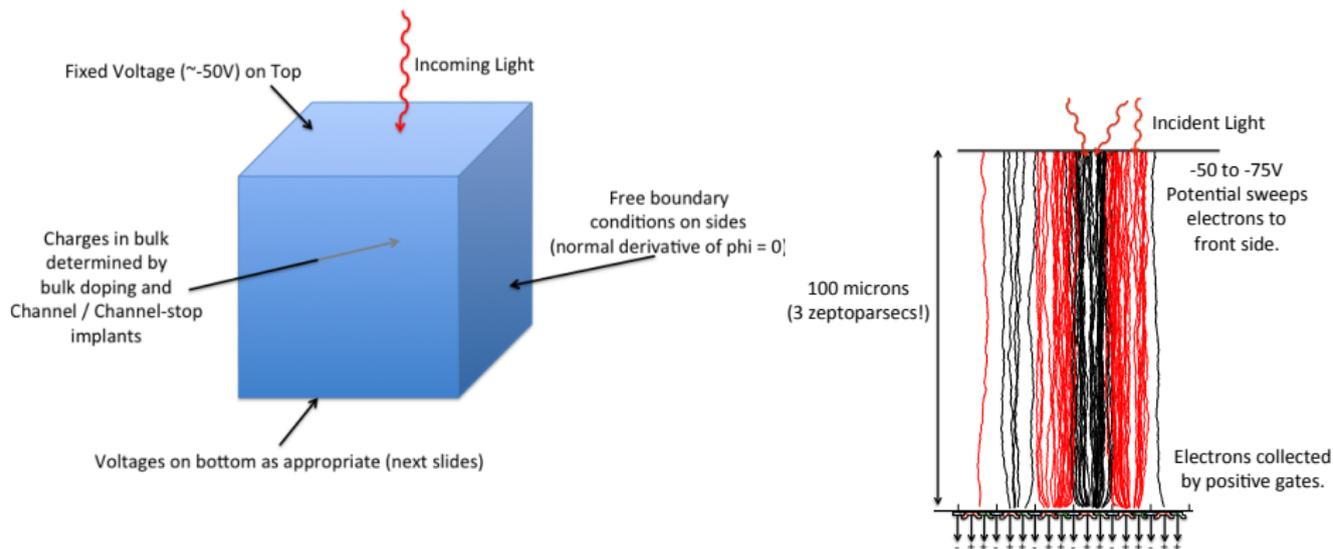
# The last 3 zeptoparsecs!



# Why build a physics-based simulation of the sensor?

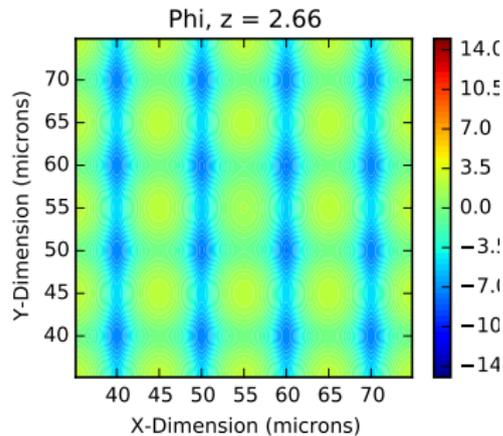
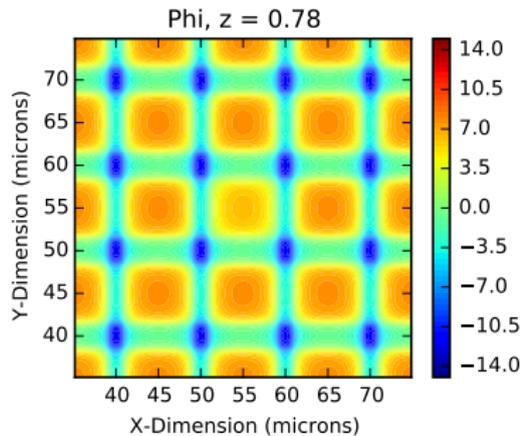
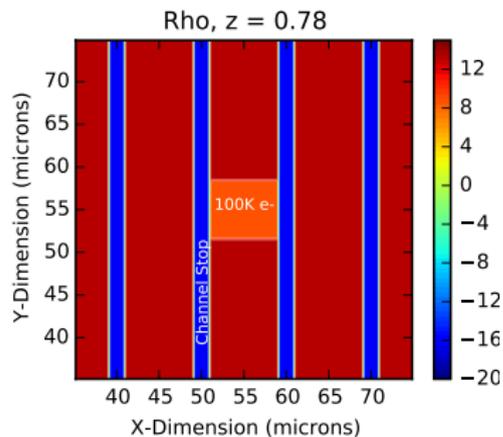
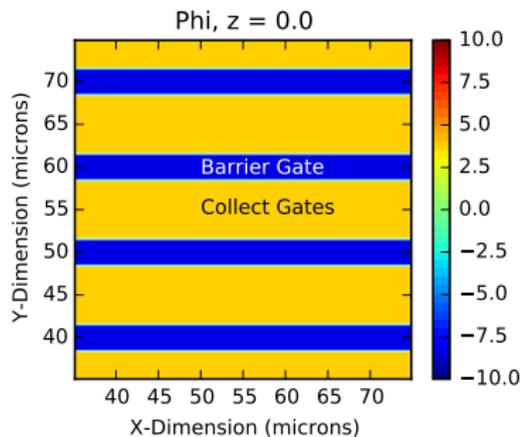
- Allows one to turn physical effects (such as diffusion) on and off.
- Allows study of small effects without interference from extraneous problems:
  - Noise
  - Crosstalk
  - Optical Distortion, ...
- Allows study of the impact of sensor differences:
  - Doping
  - Thickness
  - Vendor, ...
- Allows study of the impact of environmental effects:
  - Temperature
  - Voltage, ...

# Typical Simulation 100 $\mu\text{m}$ Cube.



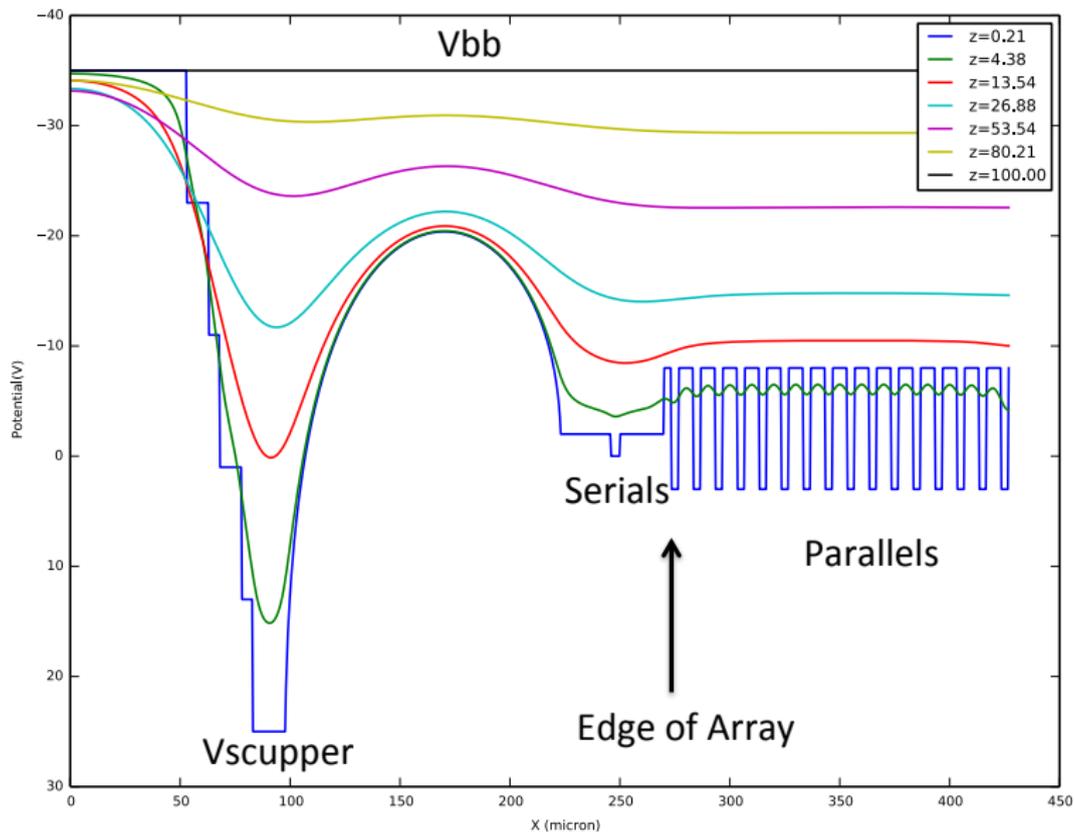
- 100 $\mu\text{m}$  Cube. - 10 X 10 pixels in X and Y.
- 32 grid cells per pixel - cell size = 0.31  $\mu$  .
- A B-F run with 256 spots, 3 million electrons ( 300,000 in central spot) takes about 6 hours.

# Pixel Array Summary Plot

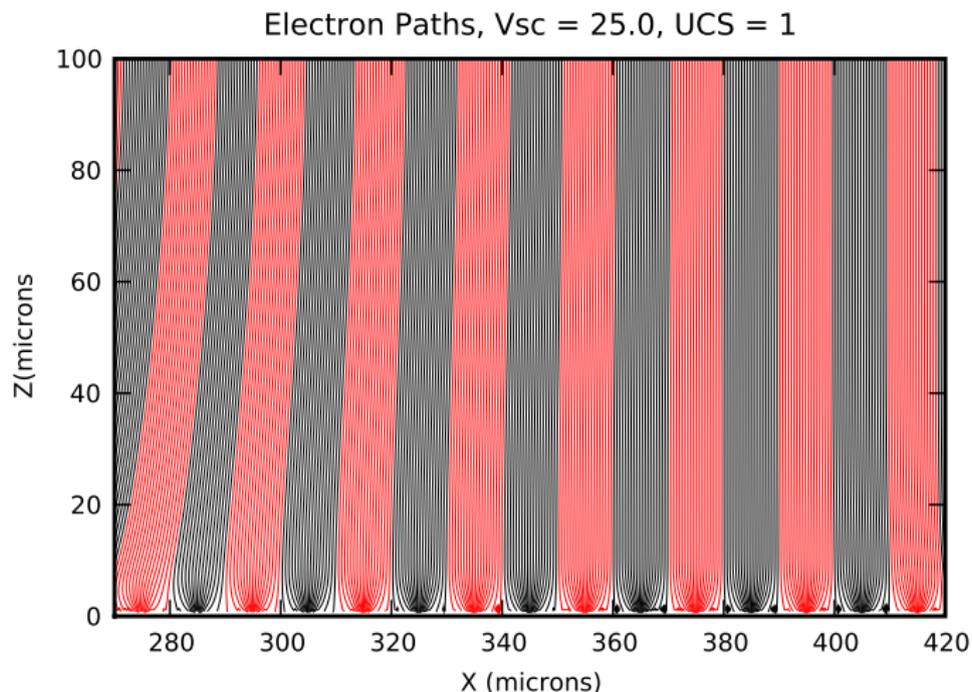


# Potential Simulation at Array Edge

Array Center Potentials. Grid = 1024\*256\*256.

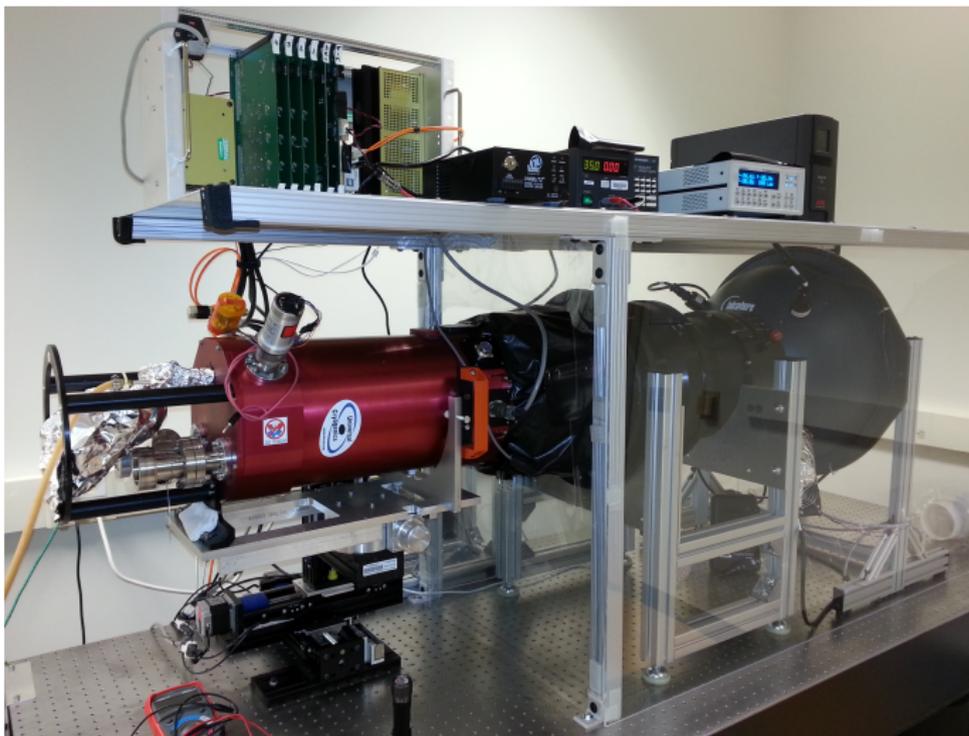


# Array Edge Electron Paths



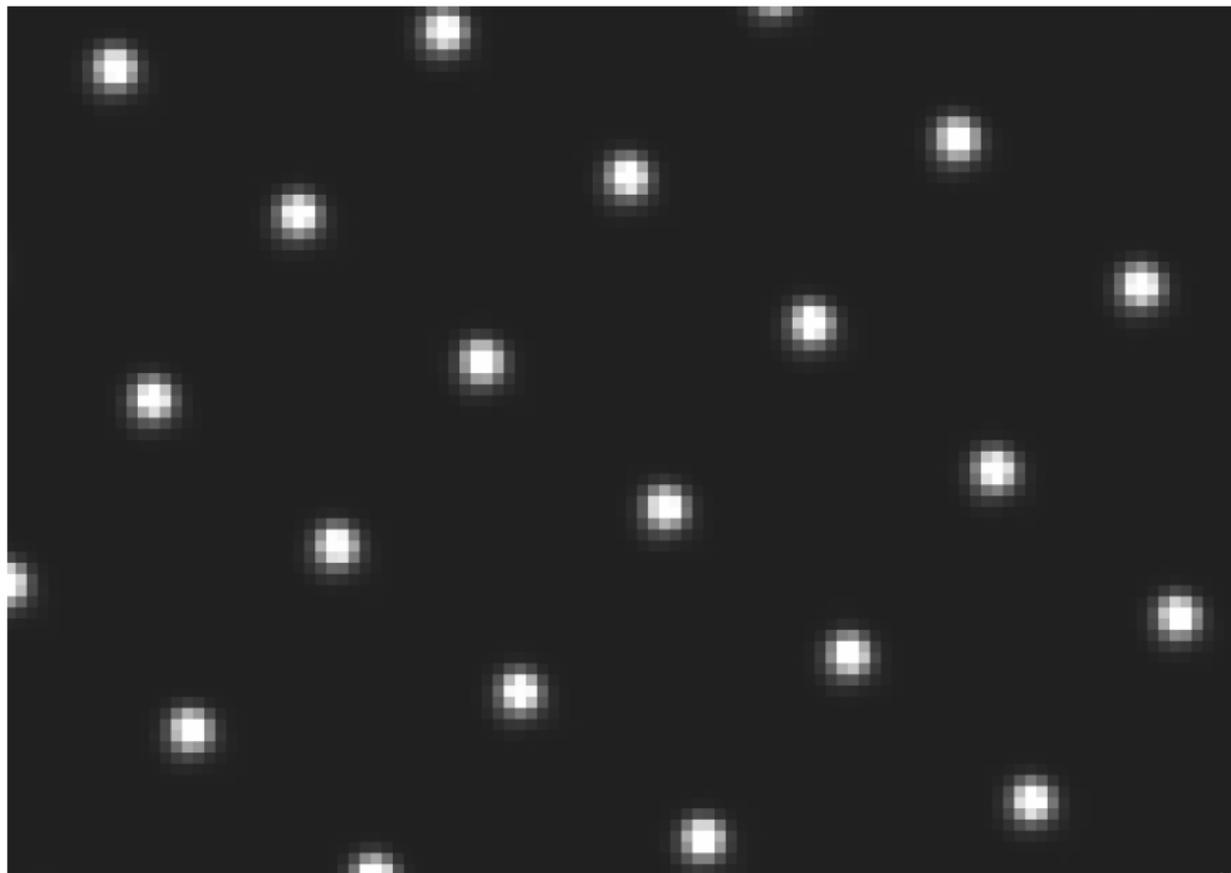
Shift at edge of array greater than 1 pixel!

# LSST Optical Simulator

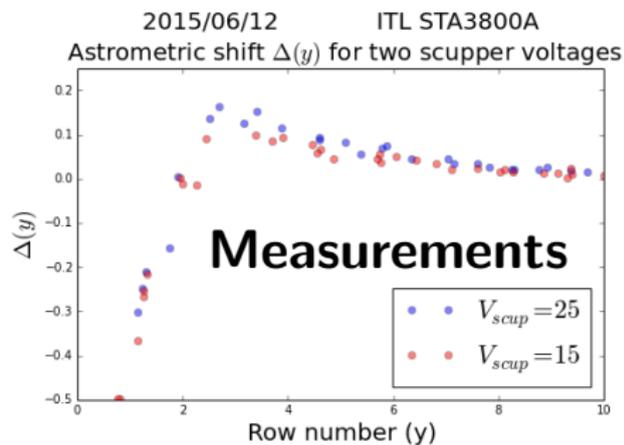
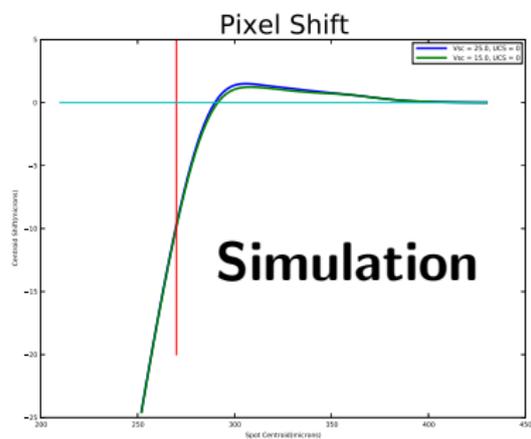


Tyson, et.al., "The LSST Beam Simulator", SPIE 9154-67 (2014),  
arXiv:1411.5667.

# Typical Image of 30 micron Spots

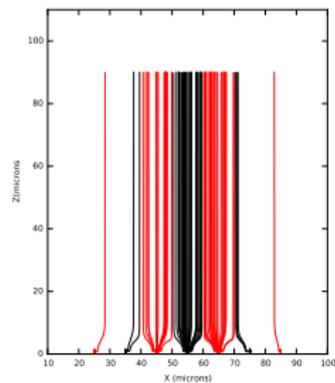
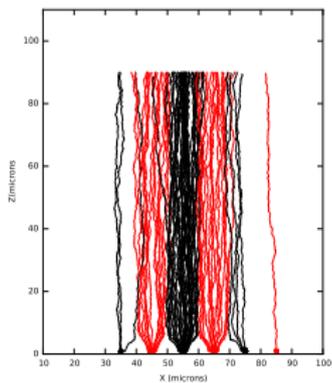
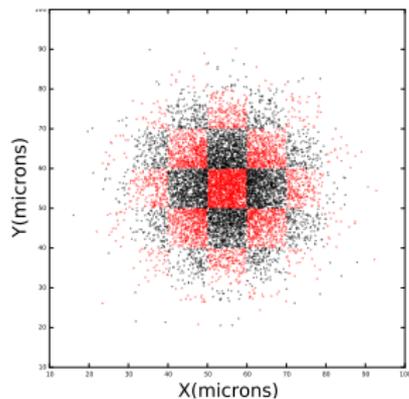
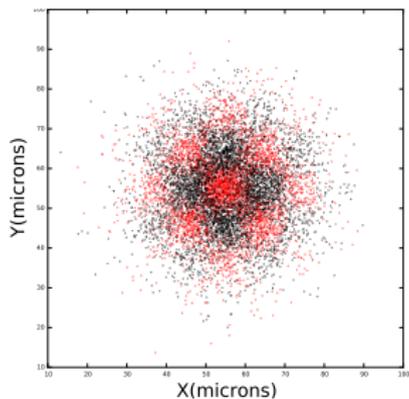


# Array Edge Astrometric Shift



Bradshaw, et al., JINST 10C4034B (2015), arXiv:1507.02683.

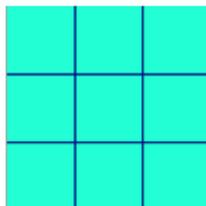
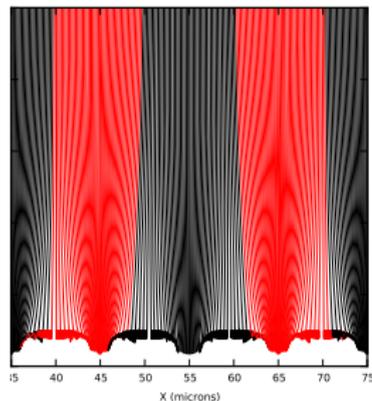
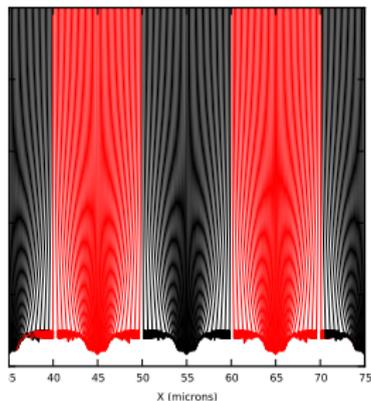
# Impact of electron diffusion



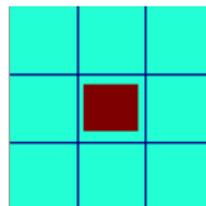
Theoretical Diffusion

Diffusion turned off

# Basics of the Brighter-Fatter Effect



Pixel empty of charge



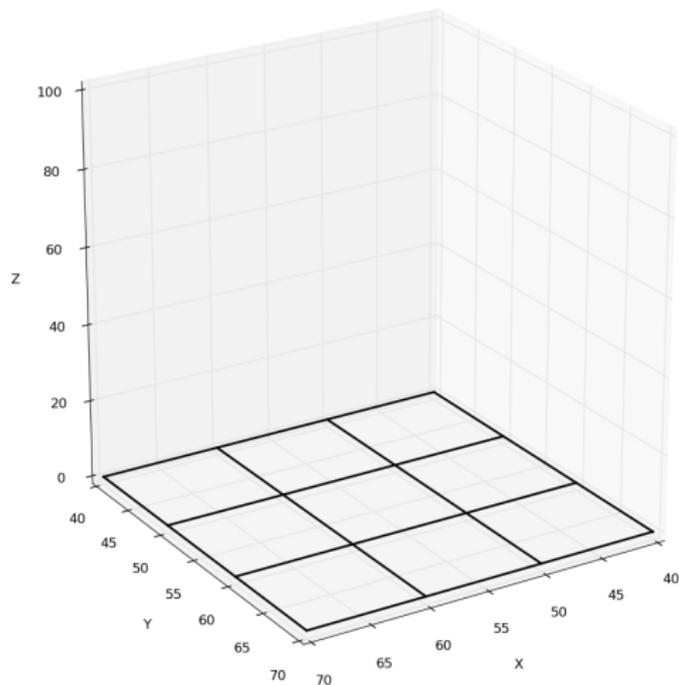
Pixel with 100K e<sup>-</sup>

- Electrons stored in the potential well repel incoming electrons and push them into surrounding pixels.

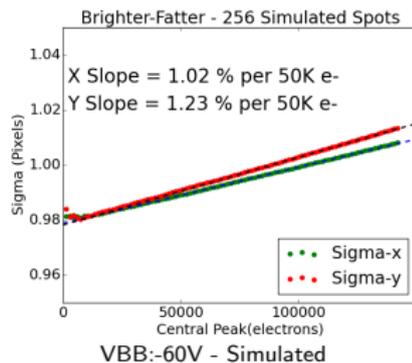
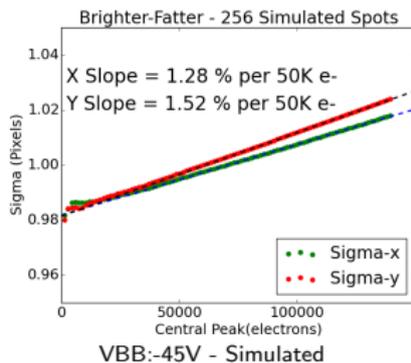
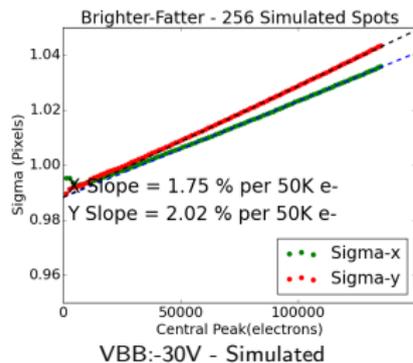
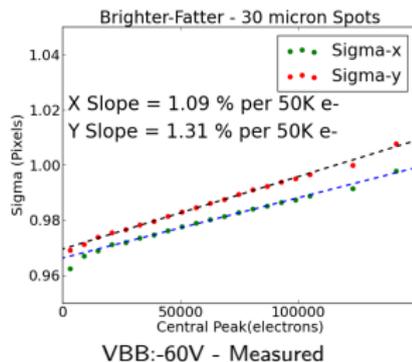
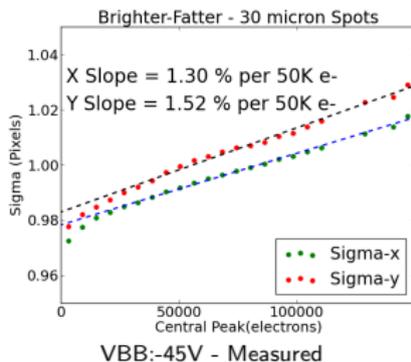
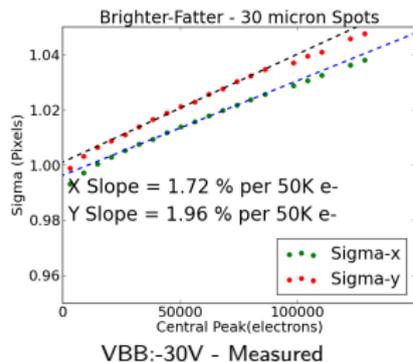
## Simulation Strategy for B-F effect.

- Solve Poisson's equation for postage stamp with all pixels empty.
- Choose a random location within the central pixel.
- Determine starting locations for  $N$  electrons in a 2D Gaussian spot.
- Propagate these electrons down to their collecting gates.
- Re-solve Poisson's equation with these wells now containing the appropriate charge.
- Repeat with  $N$  more electrons.
- I have been using 10,000 electrons per step, which places about 1000 electrons in the central pixel, so about 100 iterations are needed to fill the central pixel.
- In practice, repeat for more than one spot (typical 256), each with a different central location.

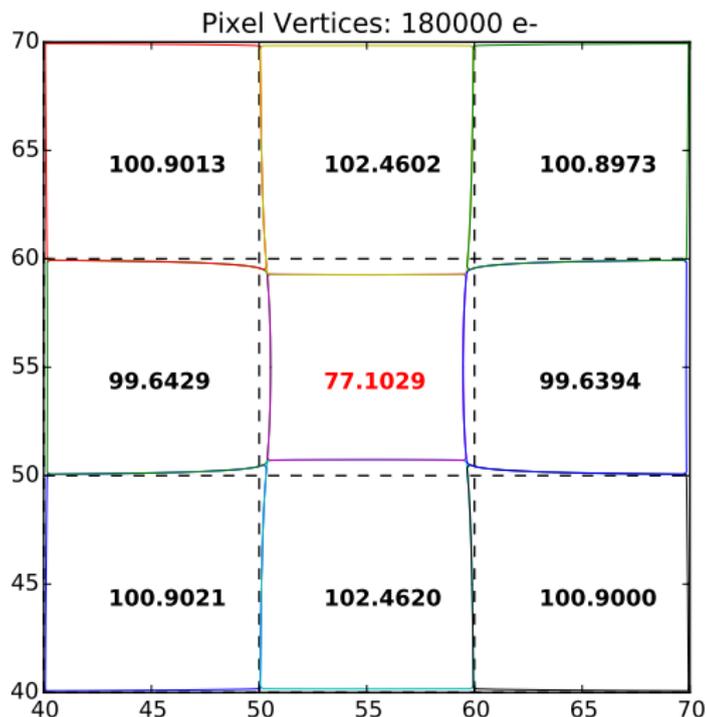
# Movie of Pixel Filling - First 10,000 Electrons



# B-F Slopes vs VBB, Measurements and Simulations - Assumed Charge Location



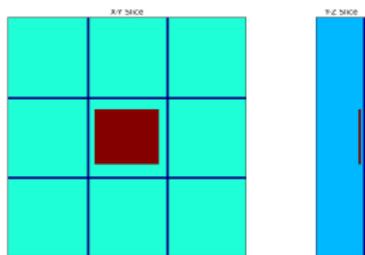
## B-F Effect and pixel correlations in CCD flats.



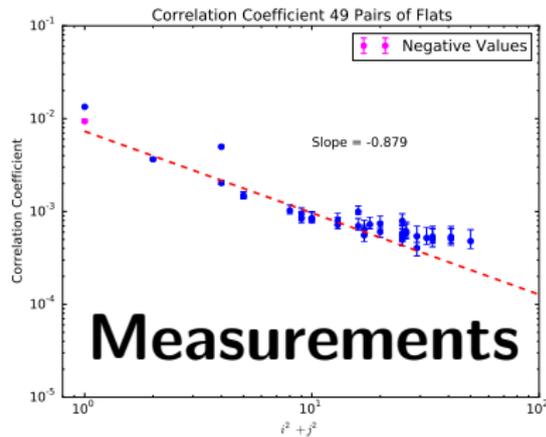
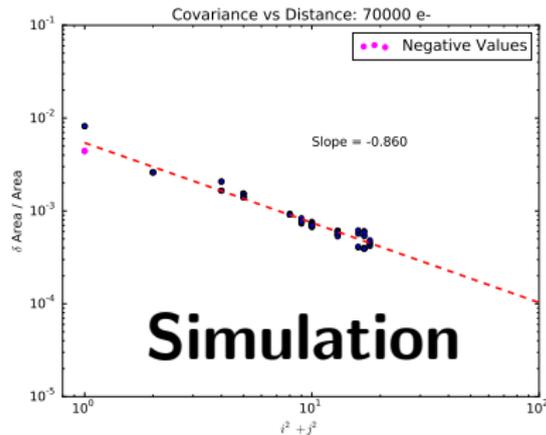
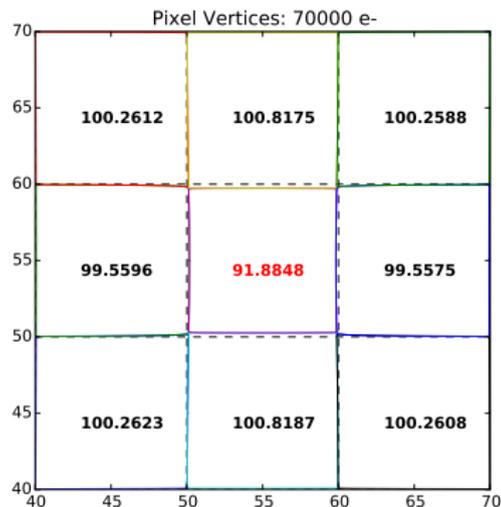
- Antilogus, et al.,  
JINST 9C3048  
(2014),  
arXiv:1402.0725.
- Rasmussen, A.,  
JINST 904027  
(2014),  
arXiv:1403.3317.

- A pixel that receives an excess of electrons causes surrounding pixels to grow slightly in area, leading them to also receive a slight excess in electrons.

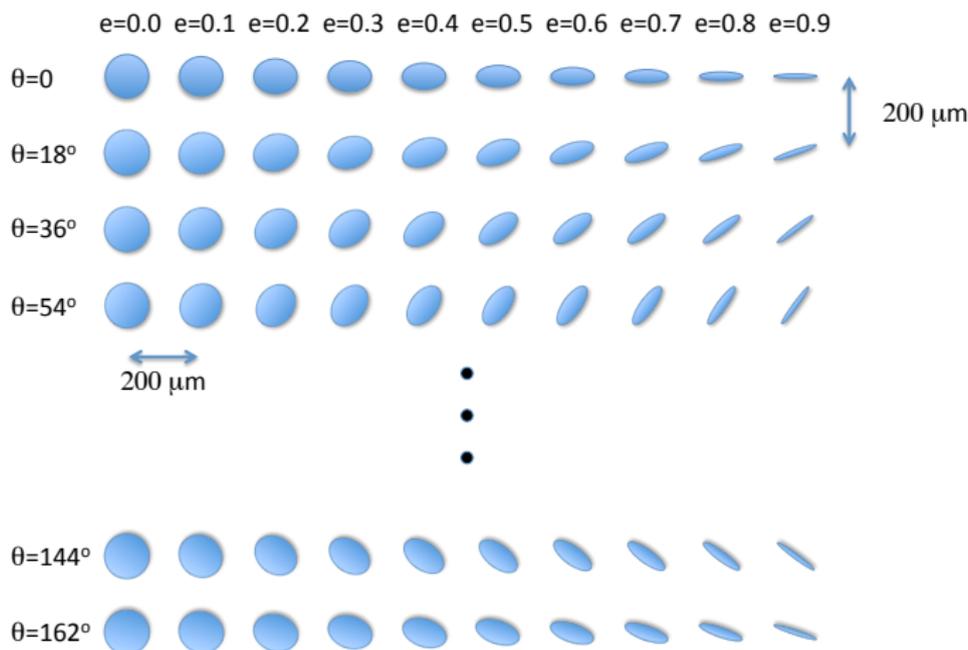
# Pixel Areas and Correlations - 70,000 e<sup>-</sup> in Central Pixel



Charge Distribution (Assumed)



# Preliminary Future Plans for Spot Masks



- A range of ellipses with different ellipticities and orientations.
- Image, extract shapes with dm stack and compare to input.
- Future masks will include mock galaxies with different light profiles.

# Conclusions and Next Steps

- Physics-based simulations are a powerful tool for studying astrometric and photometric distortions in CCDs.
- We are having good success in simulating some important effects:
  - Edge effects.
  - Brighter-Fatter effects.
  - Diffusion effects.
- Future plans include:
  - Continued refinement of the physics-based model with additional measurements:  
Voltage, Collecting phases, Saturation, ...
  - Extension to both CCD vendors.
  - More complex masks to allow tests with simulated stars and galaxies.
  - Incorporation of the learning into PhoSim.